

Chapter 14. Radiation Heat Transfer

All substances at T above 0 K emit radiation.

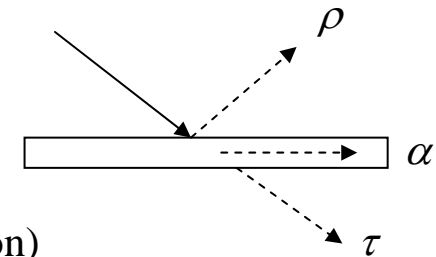
“Thermal radiation”

ρ (reflectivity: 반사율) ~ fraction of radiation falling on a body that is reflected.

α (absorptivity: 흡수율) ~ fraction that is absorbed.

τ (transmissivity: 투과율) ~ fraction that is transmitted.

$$\alpha + \rho + \tau = 1$$



Blackbody (흑체): $\alpha=1$ (a body that absorbs all incident radiation)

Emission of radiation

Wavelength of radiation: 10^{-11} cm to 10^3 m

“ important in heat flow: 0.1 to 100 μm

(visible light: 0.39 – 0.78 μm)

고온일수록 짧은 파장의 복사열을 방출

(500 °C 이상에서는 가시광선보다 큰 파장에 의한 복사도 점차 중요해 짐)

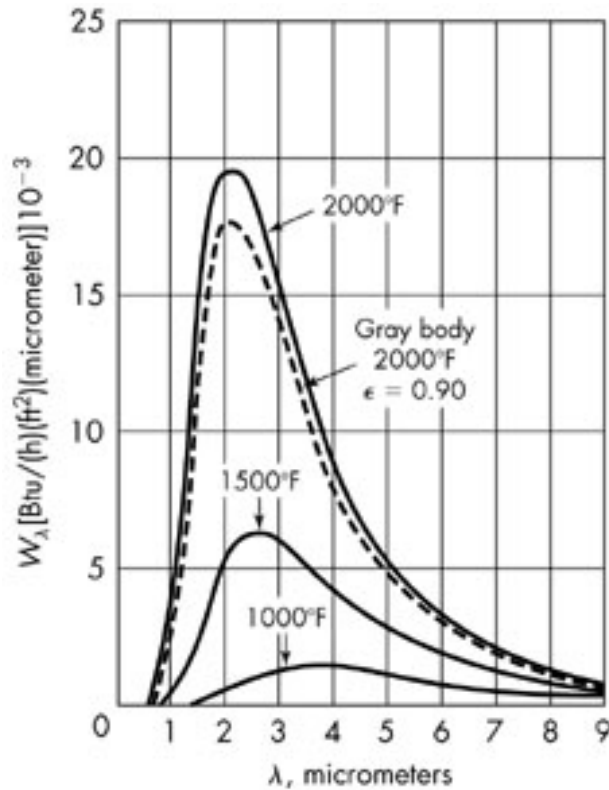


Fig. 14.1. Energy distribution in spectra of blackbodies and gray bodies.

W_λ (**monochromatic radiation power**: 단색광 복사력): Monochromatic radiation emitted from unit area in unit time divided by the wavelength

Fig. 14.1의 y축 값
(방사에너지는 온도와 파장에 의존)

radiation of a single wavelength

Total radiating power (총 복사력): $W = \int_0^\infty W_\lambda d\lambda$ → area of Fig. 14.1

$$\varepsilon \text{ (emissivity)} \equiv \frac{W}{W_b} \quad (\text{방사율})$$

emissive power of a blackbody (최대가능 복사력)

$$\varepsilon_\lambda \text{ (monochromatic emissivity)} \equiv W_\lambda / W_{b,\lambda}$$

단색광 방사율

Gray body (회색체): ε_λ of a body is the same for all λ

(단색광 방사율이 모든 파장에서 동일한 물체)

ε of solids --- increases with T .

- polished metals: 0.03 ~ 0.08
- oxidized metals: 0.6 ~ 0.85
- paper, boards: 0.65 ~ 0.95
- paints (Al계 paints 제외): 0.80 ~ 0.96

* Laws of blackbody radiation (흑체 복사 법칙)

Stefan-Boltzmann law:

$$W_b = \sigma T^4$$

Stefan-Boltzmann constant: $5.672 \times 10^{-8} \text{ W/m}^2 \cdot \text{K}^4$

Planck's law (distribution of energy in the spectrum of a blackbody):

$$W_{b,\lambda} = \frac{2\pi \mathbf{h} c^2 \lambda^{-5}}{e^{\mathbf{h}c/\mathbf{k}\lambda T} - 1} \quad \text{--- Eq. (14.6)} \quad \rightarrow \quad \text{Plots (Fig. 14.1)}$$

$$= \frac{C_1 \lambda^{-5}}{e^{C_2/\lambda T} - 1}$$

$W_{b,\lambda}$: monochromatic emissive power of blackbody
 \mathbf{h} : Plank's constant
 c : speed of light
 λ : wavelength of radiation
 \mathbf{k} : Boltzmann's constant
 C_1, C_2 : constants

Relationship between Planck's law & Stefan-Boltzmann law:

$$W_b = \int_0^{\infty} \underline{W_{b,\lambda}} d\lambda$$

Stefan-Boltzmann law

(즉, Stefan-Boltzmann 식은 Planck 식을 적분한 식)

Wien's displacement law (λ that gives the maximum $W_{b,\lambda}$):

$$T \lambda_{\max} = C$$

← Planck 식을 λ 에 대해 미분하고 이를 0으로 두어 구한 식

(T 의 단위를 K, λ_{\max} 의 단위를 μm 로 둘 경우 C 는 2,890을 사용)

Absorption of Radiation by Opaque Solids

Opaque body (불투명체)

$$\rightarrow \tau = 0, \quad \therefore \alpha + \rho = 1$$

Reflection (반사) {

- specular reflection (정반사 or 거울반사)
 - : 매끄러운 표면에서의 반사
 - 반사각은 입사각에 의존
- diffuse reflection (난반사) or diffuse radiation (난복사)
 - : 거친 표면에서의 반사
 - 일정한 반사각이 없이 모든 방향으로 반사
 - 반사율과 흡수율이 입사각에 무관
 - 보다 일반적인 경우임.

* Absorptivity of a **gray body** – the same for all wavelengths, like the emissivity
회색체의 표면이 난복사(난반사)한다면 monochromatic absorptivity도
radiant beam의 입사각에 무관

→ Total absorptivity {

- equals the monochromatic absorptivity.
- is independent of the incident angle.

Kirchhoff's law (At T equil., the ratio of the total radiating power W of any body to its absorptivity α depends only on the T of the body):

$$\frac{W_1}{\alpha_1} = \frac{W_2}{\alpha_2}$$

← total radiating powers of 2 bodies
← absorptivities of 2 bodies

If the first body is a blackbody,

$$\alpha = 1 \quad \& \quad W_1 = W_b = \frac{W_2}{\alpha_2}$$

$$\therefore \alpha_2 = \frac{W_2}{W_b} = \varepsilon_2$$

↑ *emissivity* 의 정의

→ When any body is at T equilibrium, $\varepsilon = \alpha$

. Total radiation for a unit area of an opaque body:

$$\frac{q}{A} = \sigma \varepsilon T^4$$